

Recall: we have found orthonormal 'standing' waves on the surface of sphere $\equiv Y_l^m(\theta, \phi) = |lm\rangle$

"orthonormal" $\Rightarrow \int_0^\pi \int_0^{2\pi} Y_l^m Y_l^m \sin\theta d\theta d\phi = \delta_{mm'} \delta_{ll'}$

element for "solid angle"
 $r^2 d\Omega = \text{surface area}$

$Y_l^m \sim \underbrace{f(\theta)}_{l-m \text{ zonal nodes}} e^{im\phi}$

$Y_l^0 \sim (\sin\theta e^{i\phi})^l$

$Y_l^0 \sim P_l(\cos\theta)$ ← Legendre

$[L_i, X_j] = i\hbar \epsilon_{ijk} X_k$

$[L_i, \vec{X} \cdot \vec{P}] = 0$

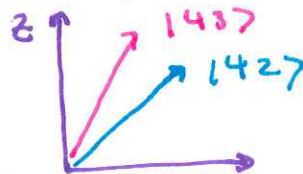
we used a subset of commutation relations to make operators that changed L_z but left L^2 unchanged

$[L^2, L_\pm] = 0$

$L_\pm = L_x \pm iL_y$

$[L_z, L_\pm] = \pm \hbar L_\pm$

$L^2 = L_+ L_- + \hbar L_z + L_z^2$



$L_\pm Y_l^m = \hbar \sqrt{l(l+1) - m(m\pm 1)} Y_l^{m\pm 1}$

$\begin{cases} L_z Y_l^m = \hbar m Y_l^m \\ L^2 Y_l^m = \hbar^2 l(l+1) Y_l^m \end{cases}$

there must be a whole number of states between the largest value of $m=l$ & the smallest $m=-l$

so $2l+1 = \text{whole number}$; $l = \text{integers or "half integers"}$

Remarks: it is easy to show via L_\pm that

$\langle L_x \rangle = \langle Y_l^m | L_x | Y_l^m \rangle = 0 \quad \& \quad \langle L_y \rangle = 0$

so the \vec{L} is usually displayed as a precessing vector on the surface of a cone so, on average, $L_x = 0$ while L_z remains constant at $m\hbar$. See Figure 4.9

$p^2 = -\hbar^2 \left(\partial_r^2 + \frac{2}{r} \partial_r \right) + \frac{L^2}{r^2}$ so using separation of variables

$\Psi = R(r) Y_l^m(\theta, \phi) \Rightarrow \left[-\frac{\hbar^2}{2m} \left(\partial_r^2 + \frac{2}{r} \partial_r \right) + \frac{\hbar^2 l(l+1)}{2m r^2} + V \right] R = ER$

if we write $R = \frac{u(r)}{r}$

alt: $\frac{1}{r} \partial_r^2 (rR)$ centrifugal potential

$\left[-\frac{\hbar^2}{2m} \partial_r^2 + V_{\text{eff}} \right] u = E u$

$V + \frac{\hbar^2 l(l+1)}{2m r^2}$

note: norm $\int \Psi^\dagger \Psi r^2 dr d\Omega$

becomes $\int u^\dagger u dr \int Y_l^m Y_l^m d\Omega$

now looks like 1d

orthonormal

Example 1: infinite spherical square well. $V(r) = \begin{cases} 0 & r < a \\ \infty & r > a \end{cases}$

Remark: the 1d square well was easy cuz the diff eq was common & produced familiar functions \sin & \cos . For higher energy states we just shortened the λ so more oscillations fit in same place. It can't be this easy in 3d as the radial functions can't be as regular as $\frac{\sin}{\cos}$ - eg as waves collapse toward origin the same stuff is concentrated in less surface so amplitude must increase - thus functions with decaying oscillations as $r \rightarrow \infty$ are the best we can hope for.

$$\left[-\frac{\hbar^2}{2m} \left(\partial_r^2 + \frac{2}{r} \partial_r \right) + \frac{\hbar^2 l(l+1)}{2m r^2} + 0 \right] R = E R$$

$\times \frac{2m}{\hbar^2}$; define $\frac{2mE}{\hbar^2} \equiv k^2$; define $\rho = kr$

$$\left(- \left(\partial_\rho^2 + \frac{2}{\rho} \partial_\rho \right) + \frac{l(l+1)}{\rho^2} \right) R = \epsilon R$$

or $\left(- \partial_\rho^2 + \frac{l(l+1)}{\rho^2} \right) u = u$ where $u = r R$

seek behavior of u or R as $\rho \rightarrow \infty \rightarrow -u'' = u$

$$u = \frac{\sin}{\cos}(\rho)$$

seek behavior of R or u as $\rho \rightarrow 0$

$$R = \frac{1}{r} \frac{\sin}{\cos}(kr)$$

Try power law: $R \sim \rho^a \Rightarrow$

$$-a(a-1)\rho^{a-2} - 2a\rho^{a-2} + l(l+1)\rho^{a-2} = \rho^a$$

$$[l(l+1) - a(a+1)]\rho^0 = \rho^2 \text{ small as } \rho \rightarrow 0$$

so $a = l$ or $-(l+1)$

Try power Law: $R = \rho^l \sum a_k \rho^k$

$l(l+1)$

$$\sum -a_k (l+k)(l+k-1) \rho^{l+k-2} - 2a_k (l+k) \rho^{l+k-2} + a_k \rho^{l+k-2} = \sum a_k \rho^{l+k}$$

shift \downarrow

$$\sum \left\{ l(l+1) - (l+k+2)(l+k+3) \right\} a_{k+2} \rho^{l+k}$$

\hookrightarrow factors: $(k+2)(2l+k+3)$

$$a_{kr2} = \frac{-1}{(k+2)(2l+k+3)} a_k \quad \text{"two term recursion"}$$

$$k \rightarrow zi : a_{i+1} = \frac{-1}{(2i+2)(2l+2i+3)} a_i = \frac{-1/4}{(i+1)(l+i+3/2)} a_i$$

$$\Rightarrow \sum a_i p^{2i} = {}_0F_1 \left(\begin{matrix} - \\ l+3/2 \end{matrix} ; -\frac{p^2}{4} \right)$$

hypergeometric: ${}_1F_1 \left(\begin{matrix} a \\ b \end{matrix} ; x \right) = \sum \frac{(a)_n}{(b)_n} \frac{x^n}{n!}$

shifted factorial or Pochhammer
 $(b)_n = b(b+1)(b+2)\dots(b+n-1)$
n terms

$$R \sim p^l {}_0F_1 \left(\begin{matrix} - \\ l+3/2 \end{matrix} ; -\frac{p^2}{4} \right) = j_l(p)$$

Boundary Conditions require $R(\varphi) \sim j_l(k\varphi) = 0$
 so $k\varphi$ must be a zero of $j_l \dots$ call B_{nl}
 so $k = \frac{B_{nl}}{a} ; E = \frac{k^2}{2na^2} B_{nl}^2$

l, n_r	root	E'
0,0	3.1	9.87
1,0	4.5	20.19
2,0	5.8	33.22
0,1	6.3	39.48
3,0	7.0	48.83
1,1	7.7	59.68
4,0	8.2	66.95
2,1	9.1	82.72
5,0	9.4	87.53
0,2	9.4	88.83

